

From Schrödinger to Boltzmann

Summary:

$$G_{11} = -i \langle T \psi(1) \psi^\dagger(2) \rangle$$

$$G_{12} = G^< = i \langle \psi^\dagger(2) \psi(1) \rangle$$

$$G_{21} = G^> = -i \langle \psi(1) \psi^\dagger(2) \rangle$$

$$G_{22} = -i \langle \tilde{T} \psi(1) \psi^\dagger(2) \rangle$$

$$G^R = -i \theta(t_1 - t_2) \langle \{ \psi(1), \psi^\dagger(2) \} \rangle$$

$$G^A = i \theta(t_2 - t_1) \langle \{ \psi(1), \psi^\dagger(2) \} \rangle$$

$$G^R - G^A = -i \langle \{ \psi(1), \psi^\dagger(2) \} \rangle$$

$$G^R - G^A = G^> - G^<$$

$$G^< = G^> + G^<$$

~~Summary~~

Equilibrium

Consider $H = \sum_p \epsilon_p c_p^\dagger c_p$

$$\psi(r, t) = \frac{1}{\sqrt{V}} \sum_p e^{i p r - i \epsilon_p t} c_p$$

ϵ_p - dispersion

V - volume

$$\psi^\dagger(r, t) = \frac{1}{\sqrt{V}} \sum_p e^{-i p r + i \epsilon_p t} c_p^\dagger$$

$$\langle c_p^\dagger c_p \rangle = f(\epsilon_p), \quad \langle c_p c_p^\dagger \rangle = 1 - f(\epsilon_p)$$

$$G^<(1, 2) = i \frac{1}{V} \sum_{p_1, p_2} e^{-i p_2 r_2 + i \epsilon_{p_2} t_2} e^{i p_1 r_1 - i \epsilon_{p_1} t_1} \underbrace{\langle c_{p_2}^\dagger c_{p_1} \rangle}_{\sim f(\epsilon_{p_1}) \delta_{p_1, p_2}}$$

$$= i \frac{1}{V} \sum_p e^{i p (r_1 - r_2) - i \epsilon_p (t_1 - t_2)} f(\epsilon_p)$$

Do Wigner transform with respect to time:

$$G^<(\epsilon, r_1, r_2) = \int G^<(1, 2) e^{i \epsilon (t_1 - t_2)} d(t_1 - t_2) =$$

$$= 2\pi i \frac{1}{V} \sum_p e^{i p (r_1 - r_2)} \delta(\epsilon - \epsilon_p) f(\epsilon),$$

Similarly

$$G^>(\epsilon, r_1, r_2) = -2\pi i \frac{1}{V} \sum_p e^{i p (r_1 - r_2)} \delta(\epsilon - \epsilon_p) (1 - f(\epsilon))$$

So, indeed, according to general relations:

$$G^> - G^< = -2\pi i \frac{1}{V} \sum_P e^{ip(\tau_1 - \tau_2)} \delta(\varepsilon - \varepsilon_p) \equiv G^R - G^A(\varepsilon, \tau_1, \tau_2)$$

you may also check it directly:

$$G_P^R = \frac{1}{\varepsilon - \varepsilon_p + i0}, \quad G_P^A = \frac{1}{\varepsilon - \varepsilon_p - i0}$$

$$\Rightarrow G_P^R - G_P^A = -2\pi i \delta(\varepsilon - \varepsilon_p)$$

$$G_{\mathbb{R}}^R(\varepsilon, \tau_1, \tau_2) - G^A(\varepsilon, \tau_1, \tau_2) = \frac{1}{V} \sum_P e^{ip(\tau_1 - \tau_2)} (-2\pi i \delta(\varepsilon - \varepsilon_p)) \leftarrow \text{checked}$$

We have proven that

$$G^<(\varepsilon, \tau_1, \tau_2) = -f(\varepsilon) (G^R - G^A)$$

$$f(\varepsilon) = \frac{1}{e^{\frac{\varepsilon - \mu}{2T}} + 1}$$

$$G^> = (1 - f(\varepsilon)) (G^R - G^A)$$

$$\Rightarrow G^K = G^> + G^< = \underbrace{(1 - 2f(\varepsilon))}_{\tanh \frac{\varepsilon - \mu}{2T}} (G^R - G^A)$$

Towards Boltzmann Equation

We now separate the Hamiltonian to a non-interacting and interacting parts:

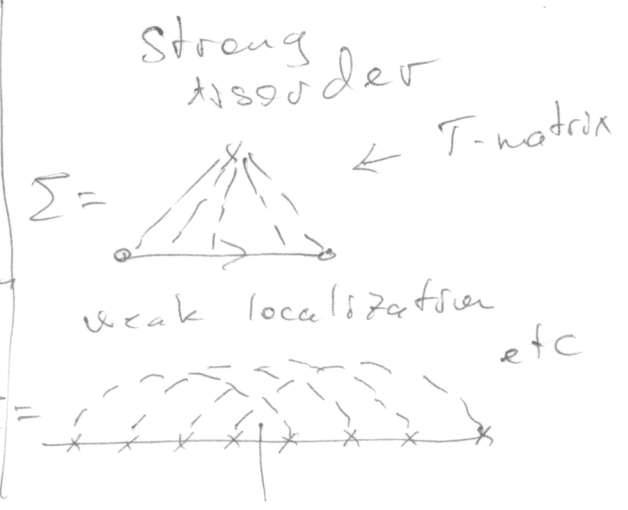
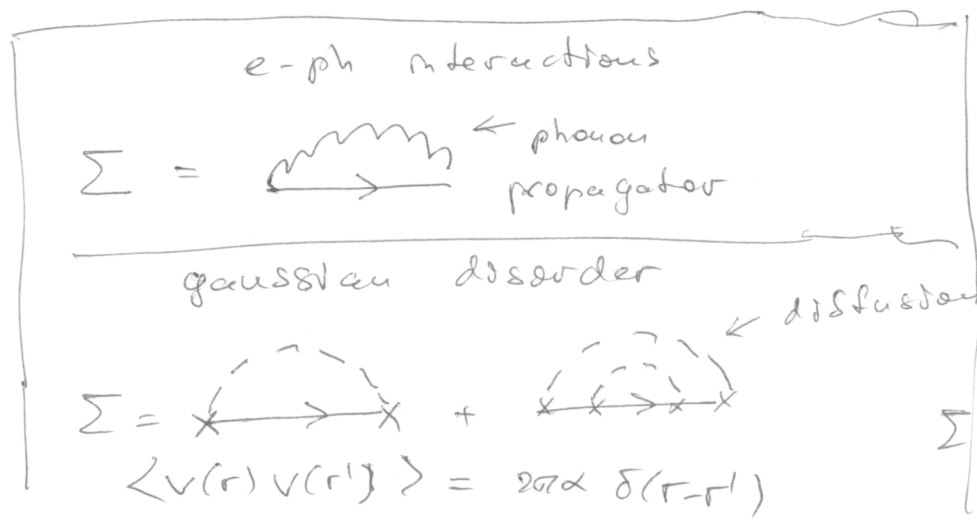
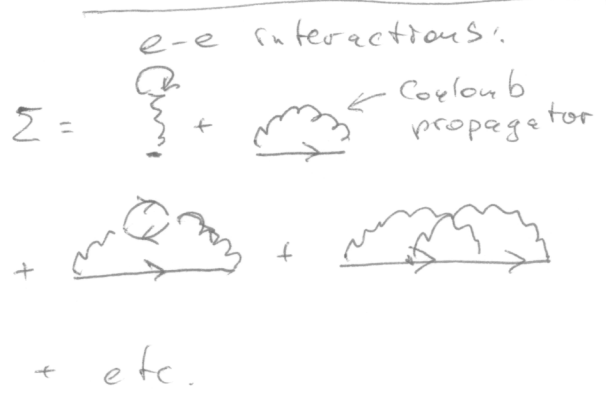
$$H \Rightarrow H_0 + H_{int}$$

↑
eigenstates can be found exactly

Quasiparticles that are described by the non-interacting part H_0 will now acquire a finite life-time due to interactions.

This process is described in some desired approximation by the object called Self-energy Σ

Interactions can be anything: ^{Coulomb} interactions between electrons, electron-phonon interactions, scattering of electrons on disorder potential, etc.



Self-energy modifies G_0 in such a way that quasiparticles acquire a finite life time (spectrum also can change) of course

$G = G_0 + G_0 \circ \Sigma \circ G$ ← new Dyson equation

~~$G_0^{-1} \circ (G - \Sigma \circ G) = \mathbb{1}$~~

$G_0^{-1} G = \mathbb{1} + \Sigma \circ G$ ← convolution in space and time

$G_0^{-1} \equiv i\partial_t - \mathbb{H}$, $\mathbb{1} = \delta(1,2)$ — in real space

We have seen already that $\hat{G}_0 = \begin{pmatrix} G_0^R & G_0^K \\ 0 & G_0^A \end{pmatrix}$ satisfies

$\begin{pmatrix} i\partial_{t_1} - \mathbb{H}_1^{(0)} + i0 & 0 \\ 0 & i\partial_{t_1} - \mathbb{H}_1^{(0)} - i0 \end{pmatrix} \begin{pmatrix} G_0^R(1,2) & G_0^K(1,2) \\ 0 & G_0^A(1,2) \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \delta(1,2)$

\hat{G}_0^{-1} — in operator sense

Now we have the following operator identities:

$$G_0^{-1} G = \hat{1} + \Sigma \circ G$$

$$G G_0^{-1} = \hat{1} + G \circ \Sigma$$

$$\Sigma = \begin{pmatrix} \Sigma^R & \Sigma^K \\ 0 & \Sigma^A \end{pmatrix}$$

subtracting them we find

$$G_0^{-1} G - G G_0^{-1} = \Sigma \circ G - G \circ \Sigma$$

← Keldysh component
(or more precisely
 $\frac{1}{2}(K - R + A)$ - component)

Boltzmann Eq. is the equation on $\frac{f(\epsilon, p, R, T)}{\text{Wigner-transform}}$

of this equation is the Boltzmann equation!

G^K is often parameterized as

$$f = \frac{1-h}{2}$$

$$G^K = G^R \circ h - h \circ G^A \quad \leftarrow \text{convolution}$$

$$(G^K = \frac{1}{2}(G^K - G^R + G^A))$$

$$\Rightarrow G^K = - G^R \circ f + f \circ G^A$$

f - non-equilibrium distribution function!

Before we consider examples we have to do one basic exercise that is central for treating the operator equation:

$$\underline{G_0^{-1}G - GG_0^{-1} = \Sigma \circ G - G \circ \Sigma}$$

We have to write this equation in Wigner coordinates: ϵ, p, R, t , since Boltzmann equation is the equation in Wigner coordinates!

In order to do that we have to know what is the Wigner transform of a convolution $A \circ B$ and how does it can it be expressed through the Wigner transforms of A & B individually. Example: (consider only time)

$$A \circ B(t_1, t_2) = \int_{-\infty}^{+\infty} dt_3 A(t_1, t_3) B(t_3, t_2)$$

$$\text{we have: } A(t_1, t_2) = \int_{-\infty}^{+\infty} \frac{d\epsilon}{2\pi} e^{-i\epsilon(t_1 - t_2)} \underbrace{A(\epsilon, \frac{t_1 + t_2}{2})}_{\text{Wigner transform}}$$

How to express $A \circ B(\epsilon, t)$ through $A(\epsilon, t)$ & $B(\epsilon, t)$?

introduce

$$\tau = t_1 - t_2$$

$$t = \frac{t_1 + t_2}{2}$$

$$\Rightarrow (A \circ B)(\epsilon, t) = \int_{-\infty}^{+\infty} d\tau e^{i\epsilon\tau} \int_{-\infty}^{+\infty} dt_3 A\left(t + \frac{\tau}{2}, t_3\right) B\left(t_3, t - \frac{\tau}{2}\right)$$

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$$t_1 = t + \frac{\tau}{2}$$

$$t_2 = t - \frac{\tau}{2}$$

$$= \int d\tau e^{i\epsilon\tau} \int dt_3 \int \frac{d\epsilon_1}{2\pi} \int \frac{d\epsilon_2}{2\pi} e^{-i\epsilon_1 \left(\frac{t + \frac{\tau}{2} - t_3}{t_1 - t_3} \right)} e^{-i\epsilon_2 \left(\frac{t_3 - t + \frac{\tau}{2}}{t_3 - t_2} \right)} A\left(\epsilon_1, \frac{t + \frac{\tau}{2} + t_3}{2}\right) B\left(\epsilon_2, \frac{t - \frac{\tau}{2} + t_3}{2}\right)$$

Use $A\left(\epsilon_1, \frac{t_1 + t_3}{2}\right) = A\left(\epsilon_1, \underbrace{\frac{t_1 + t_2}{2}}_t + \frac{t_3 - t_2}{2}\right) = e^{\frac{t_3 - t_2}{2} \frac{\partial}{\partial t}} A(\epsilon_1, t)$

Similarly $B\left(\epsilon_2, \frac{t_3 + t_2}{2}\right) = B\left(\epsilon_2, \frac{t_1 + t_2}{2} + \frac{t_3 - t_1}{2}\right) = e^{\frac{t_3 - t_1}{2} \frac{\partial}{\partial t}} B(\epsilon_2, t)$

$$\Rightarrow (A \circ B)(\epsilon, t) = \int d\tau e^{i\epsilon\tau} \int dt_3 \int \frac{d\epsilon_1}{2\pi} \int \frac{d\epsilon_2}{2\pi} e^{-i\epsilon_1(t_1 - t_3)} e^{-i\epsilon_2(t_3 - t_2)} \left(e^{\frac{t_3 - t_2}{2} \frac{\partial}{\partial t}} A(\epsilon_1, t) \right) \left(e^{\frac{t_3 - t_1}{2} \frac{\partial}{\partial t}} B(\epsilon_2, t) \right)$$

powers of $(t_3 - t_2)$ can be obtained by differentiating the exponent $e^{-i\epsilon_2(t_3 - t_2)}$ over energy ϵ_2

These derivatives can be then passed by integration by parts on the function $B(\epsilon_2, t)$ (and same with A)

$\Rightarrow \frac{t_3 - t_2}{2} \Rightarrow -\frac{i}{2} \frac{\partial}{\partial \varepsilon_2} \leftarrow \text{acting on B function}$

$\frac{t_3 - t_1}{2} \Rightarrow \frac{i}{2} \frac{\partial}{\partial \varepsilon_1} \leftarrow \text{acting on A function}$

$\Rightarrow \tau = t_1 - t_2$

$$(A \circ B)(\varepsilon, t) = \int d\varepsilon e^{i\varepsilon\tau} \int dt_3 \int \frac{d\varepsilon_1}{2\pi} \int \frac{d\varepsilon_2}{2\pi} e^{-i\varepsilon_1(t_1 - t_3)} e^{-i\varepsilon_2(t_3 - t_2)} \left(e^{-\frac{i}{2} \frac{\partial}{\partial \varepsilon_2} \frac{\partial}{\partial t}} \left(e^{\frac{i}{2} \frac{\partial}{\partial \varepsilon_1} \frac{\partial}{\partial t}} A B \right) \right)$$

\swarrow acting on A \nwarrow acting on B
 A B

gives $2\pi \delta(\varepsilon_1 - \varepsilon_2)$ $t_1 = t + \frac{\tau}{2}, t_2 = t - \frac{\tau}{2}$

gives $2\pi \delta(\varepsilon - \frac{\varepsilon_1 + \varepsilon_2}{2})$

Eventually

$$(A \circ B)(s, t) = e^{\frac{i}{2} \left(\frac{\partial^A}{\partial \varepsilon} \frac{\partial^B}{\partial t} - \frac{\partial^A}{\partial t} \frac{\partial^B}{\partial \varepsilon} \right)} A(\varepsilon, t) B(\varepsilon, t)$$

\swarrow acting on A \nwarrow acting on B \swarrow acting on t \nwarrow acting on B

Gradient expansion

(11)

$A(\varepsilon, t), B(\varepsilon, t)$ - smooth functions of both ε , and t

$$\Rightarrow (A \circ B)(\varepsilon, t) = A(\varepsilon, t) B(\varepsilon, t) + \underbrace{\frac{i}{2} \left(\frac{\partial A}{\partial \varepsilon} \frac{\partial B}{\partial t} - \frac{\partial A}{\partial t} \frac{\partial B}{\partial \varepsilon} \right)}_{\text{Poisson bracket}} + \dots$$

Same for the Wigner transform in (r_1, r_2)

$$(A \circ B)(R, p) = A(R, p) B(R, p) - \frac{i}{2} \left(\frac{\partial A}{\partial p} \frac{\partial B}{\partial R} - \frac{\partial A}{\partial R} \frac{\partial B}{\partial p} \right)$$

different sign due to convention

Now we shall return back to our Boltzmann equation:

$$\underbrace{G_0^{-1} G - G G_0^{-1} = \Sigma \circ G - G \circ \Sigma}$$

$$G_0^{-1} G - G G_0^{-1} = [\Sigma, G]$$

Let us first consider the left-hand part:

inverse direction by integration by parts

$$G_0^{-1} G - G G_0^{-1} = (i\partial_{t_1} - H_1) G(t_1, t_2) - G(t_1, t_2) (-i\partial_{t_2} - H_2)$$

go to Wigner coordinates: $t = \frac{t_1 + t_2}{2}$, $\tau = t_1 - t_2$

$$\frac{\partial}{\partial t_1} = \frac{1}{2} \frac{\partial}{\partial t} + \frac{\partial}{\partial \tau} \Rightarrow \frac{\partial}{\partial t_1} + \frac{\partial}{\partial t_2} = \frac{\partial}{\partial t}$$

$$\frac{\partial}{\partial t_2} = \frac{1}{2} \frac{\partial}{\partial t} - \frac{\partial}{\partial \tau}$$

unit matrix in Keldysh

$$H(\epsilon, p, R, t) \sim \mathbb{1}$$

Consider $[H, G] \cong i (\nabla_R H \nabla_p G - \nabla_p H \nabla_R G)$

in gradient expansion

(here we take the leading contribution with respect to ϵ, t)

Wigner transform of H is just a classical

Hamiltonian like: $H = \frac{p^2}{2m} + U(R, t)$ ← external perturbation!

$$\nabla_R H = \nabla_R U(R) = -\vec{F} \leftarrow \text{force}$$

$$\nabla_p H = \vec{v} \leftarrow \text{velocity}$$

$$\Rightarrow [H, G] = -i (\vec{F} \vec{\nabla}_p + \vec{v} \cdot \nabla_R) G$$

$$\Rightarrow G_0^{-1} G - G G_0^{-1} \cong i (\partial_t + \vec{F} \vec{\nabla}_p + \vec{v} \cdot \nabla_R) G$$

leading order in gradients

typical Boltzmann operator

thus we

For $[\Sigma, G]$ we take the leading term in the gradient expansion, i.e. replace the Wigner transform of the product by the product of Wigner transforms:

$$\Rightarrow \underbrace{(\partial_t + \bar{F} \bar{\nabla}_p + \bar{\partial} \cdot \bar{\nabla}_R)}_{\hat{G} \text{ - notation}} \hat{G}(\varepsilon, t, R, p) \Rightarrow i [\hat{\Sigma}, \hat{G}]$$

All matrices in rotated Keldysh space

$$G^R = \frac{1}{\varepsilon - \varepsilon_p - U(R, t) + i0} ; \left(F = -\nabla_R U \right)$$

external force!

in components: \Rightarrow

$$\hat{\downarrow} G^R = 0$$

$$\hat{\downarrow} G^A = 0$$

$$\hat{\downarrow} G^K = -i \left(\underbrace{\Sigma^R G^K + \Sigma^K G^A - G^R \Sigma^K - G^K \Sigma^A}_{\text{these are how usual products of Wigner transforms}} \right)$$

$$\Rightarrow \hat{\downarrow} G^K = i \left(\Sigma^K (G^R - G^A) - G^K (\Sigma^R - \Sigma^A) \right)$$

Let us now use: the definition of the distribution function:

$$G^k = G^R \circ h - h \circ G^A$$

In Wigner transform (leading order):

$$G^k = (G^R - G^A) h \quad \text{note that } \mathcal{L}G^{R,A} = 0$$

$$\Rightarrow (G^R - G^A) \hat{\mathcal{L}}h = i(\Sigma^k - h(\Sigma^R - \Sigma^A)) (G^R - G^A) \quad (*)$$

It looks like we can cancel $G^R - G^A$, but remember

that in equilibrium $\underbrace{G^R - G^A = -2\pi i \delta(\epsilon - \epsilon_p \mp U)}$

Therefore Eq. (*) has little or no sense outside

the mass shell $\epsilon = \epsilon_p$ (if U is small !!!)

We shall integrate Eq. (*) over ϵ instead:

$$\Rightarrow \left(\frac{\partial}{\partial t} + \vec{F} \cdot \vec{\nabla}_p + \vec{\nabla} \cdot \vec{\nabla}_R \right) h_p = i \left(\Sigma^K(\epsilon_p, p, R, t) - h_p (\Sigma^R - \Sigma^A) \right)$$

Introduce $h_p(R, t) = -\frac{1}{2\pi i} \int d\epsilon G^K(\epsilon, p, R, t)$

(Note that $\vec{v} = \vec{\nabla}_p \epsilon_p$ - group velocity)

Classical Boltzmann equation!

Note that Σ^K depends on h_p in the integral way, so the equation is actually integro-differential

In equilibrium one must find $\Sigma^K = h (\Sigma^R - \Sigma^A)$ that is also a test for your calculation of Σ

Exercise! Do a bit better by considering better approximation for $[\Sigma, G]$ that includes Poisson bracket term, Take all terms with the gradients to the left-hand side of the equation!

Most of success of this approach goes in the following

directions:

- 1) Situations when \hat{H} and \hat{G} have ^{additional} matrix structure apart from Keldysh
 - a) due to spin (spin-orbit, magnetism)
 - b) due to superconductivity (electron-hole basis)
 - c) due to pseudo-spin (valleys + sublattices, Kane-Mele models)

2) ~~Qegr~~ Quasichlassical approximation

$$\hat{g}(p, R, t) = \frac{i}{\pi} \int_{-\infty}^{+\infty} d\varepsilon \hat{G}(\varepsilon, p, R, t)$$

← quasichlassical Green's function!



for a metal $g^R = 1, g^A = -1 \Rightarrow \hat{g} = \begin{pmatrix} 1 & 2h \\ 0 & -1 \end{pmatrix}$

note that $\hat{g}^2 = 1$

for superconductor: \hat{g} is a bit more complex

(see Usadel equation)

but the ~~non~~-linear constraint $\hat{g}^2 = 1$ remains

This defines non-linear theory of non-equilibrium superconductivity !!